

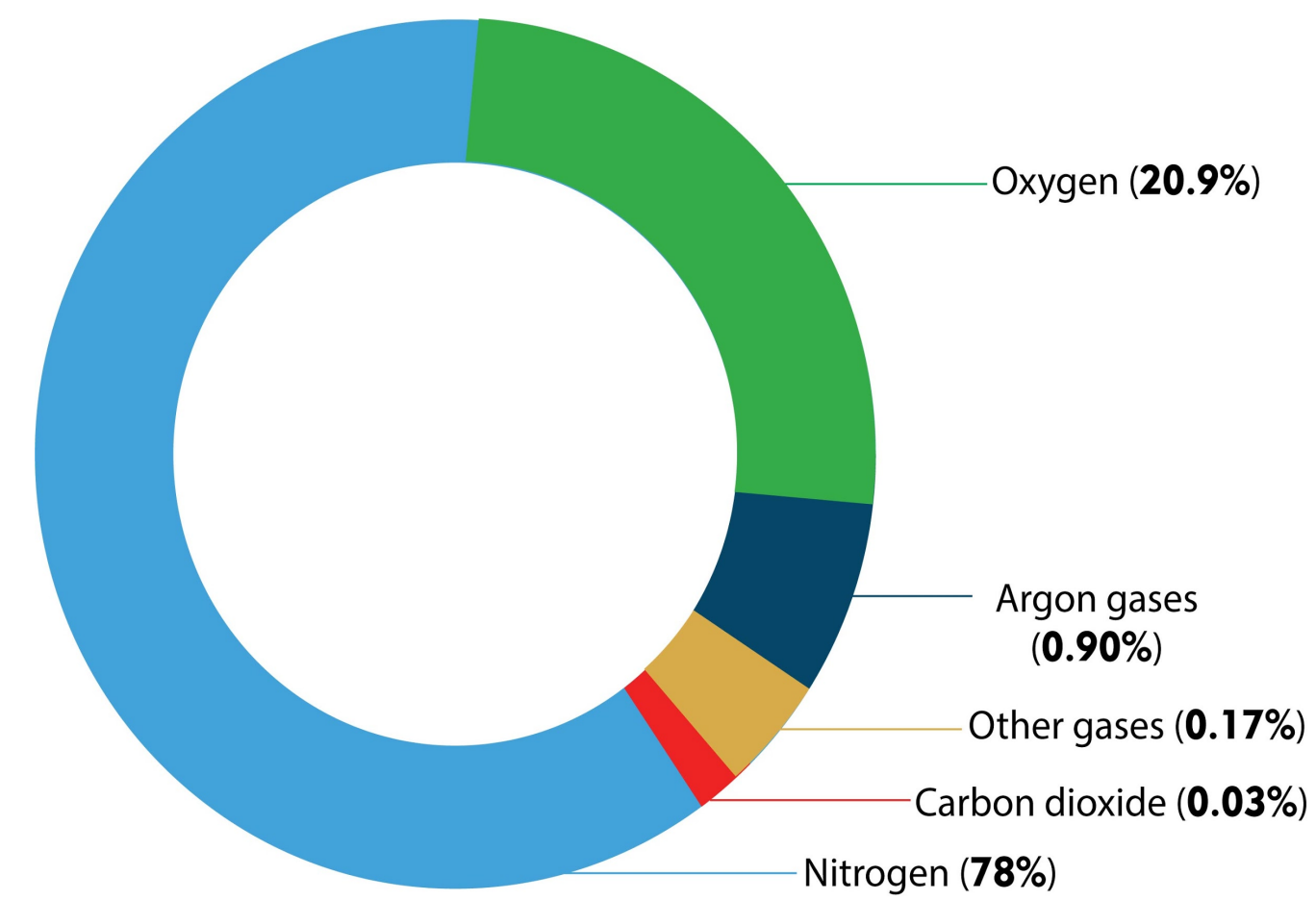
The Global Oxygen Cycle

Presentation on “Geobiology”

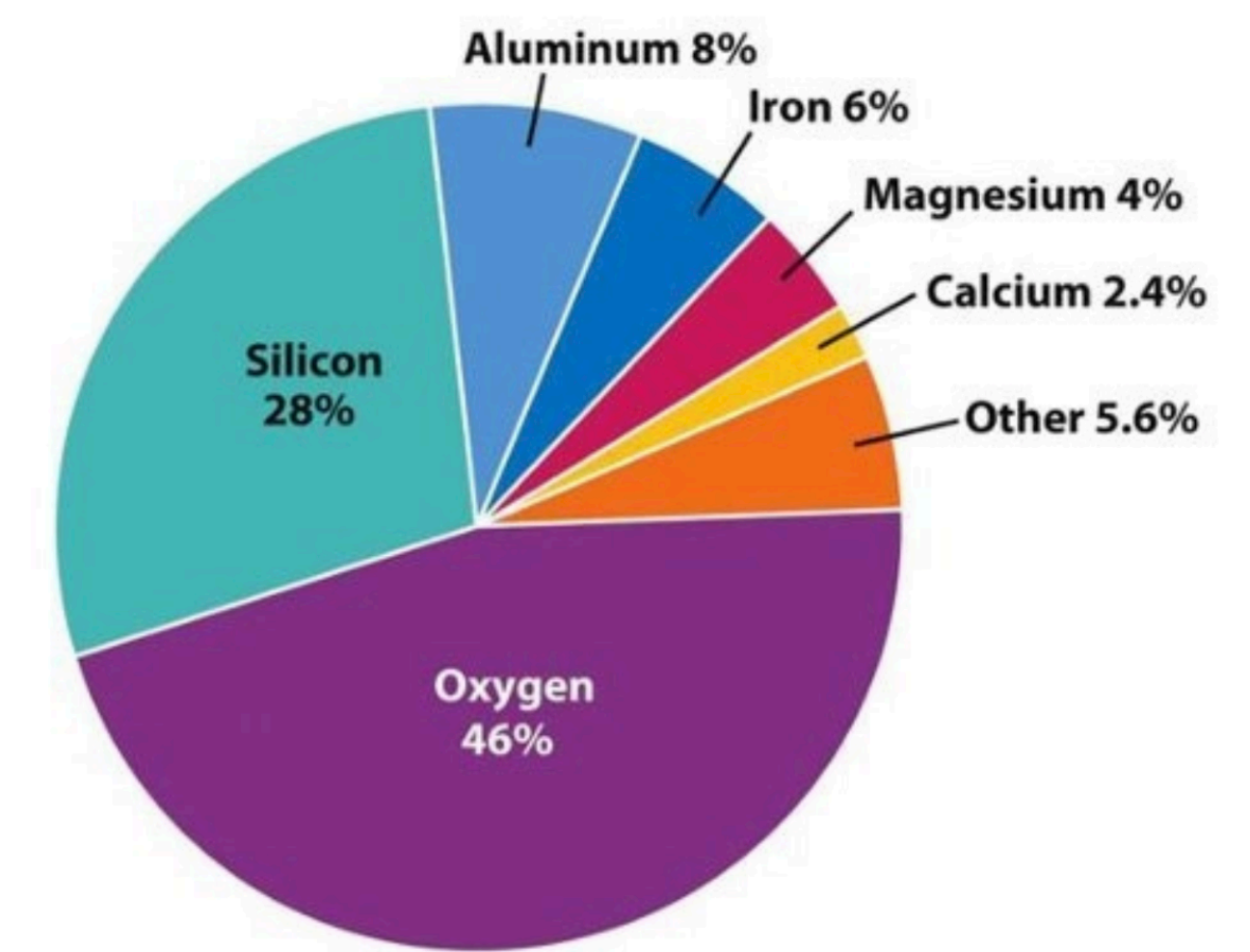
Boding Ouyang

Introduction

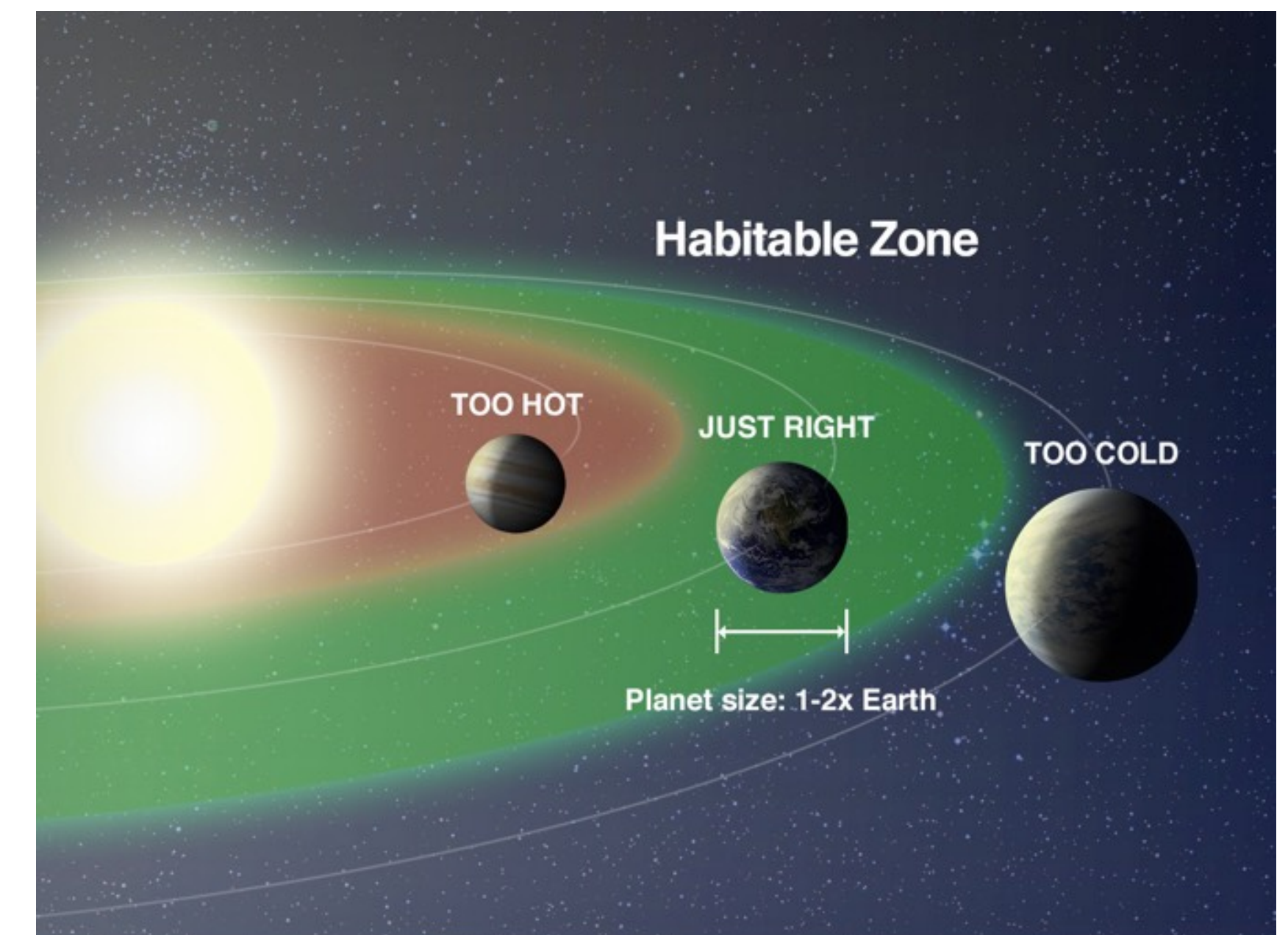
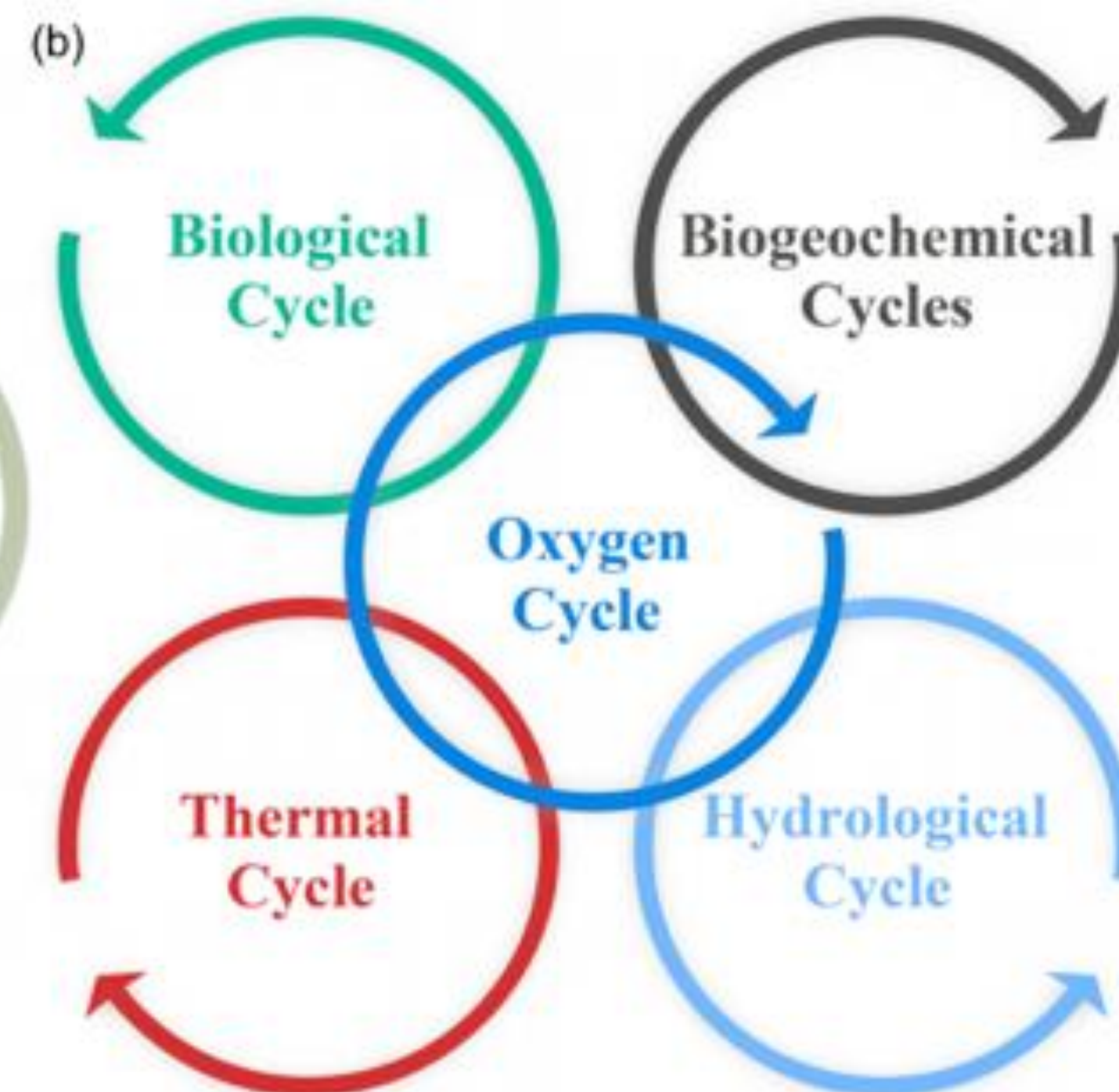
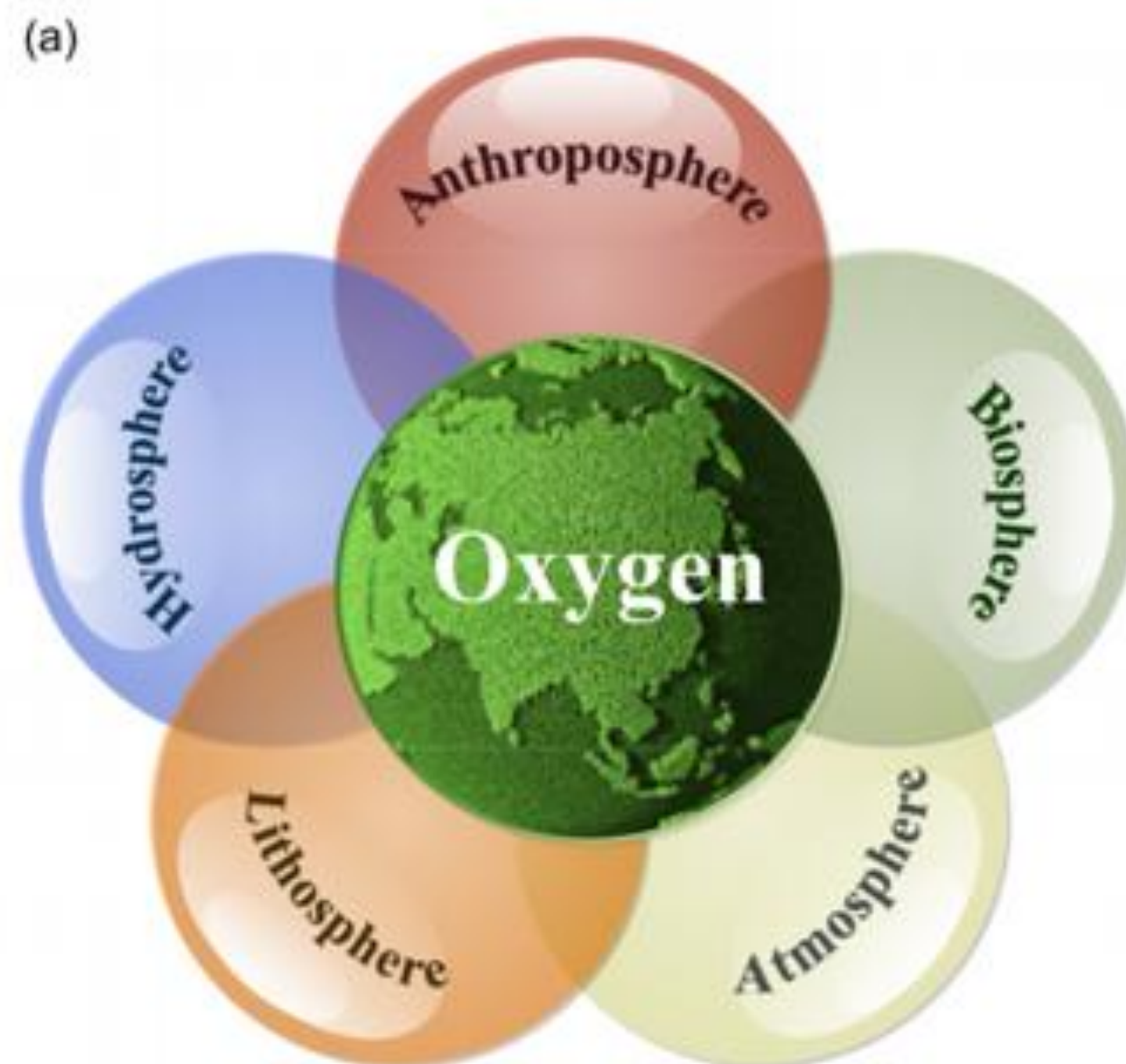
Why Study Oxygen?



Atmosphere



Crust



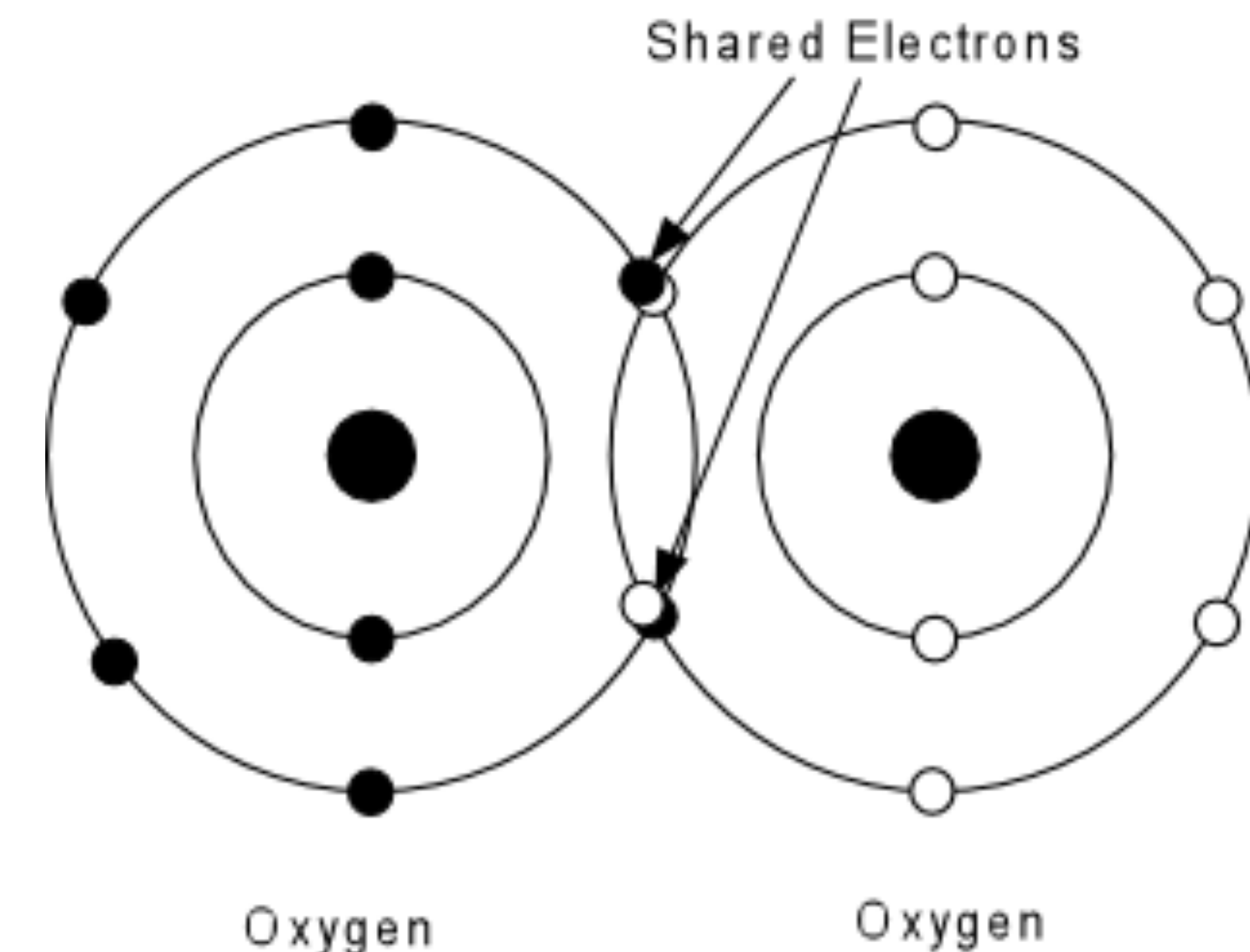
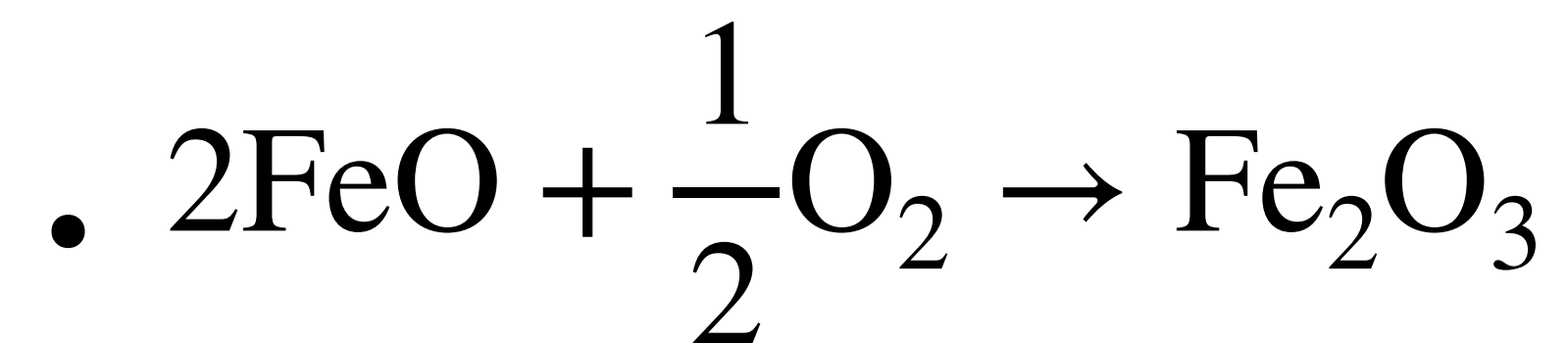
Habitability?

Chemistry & Biochemistry of Oxygen

Electron Acceptor

- Oxygen has six valence electrons, needs two more for a stable octet
- Molecular oxygen stably exists as O₂, each atom shares two electrons

- Highly electronegative, a strong electron acceptor
- Eg. Oxidation of ferrous iron (Fe²⁺) to ferric iron (Fe³⁺)



Chemistry & Biochemistry of Oxygen

Biological Role of O₂

- Respiration / Decay
 - $\text{CH}_2\text{O} + \text{O}_2 \rightarrow \text{CO}_2 + \text{H}_2\text{O}$
 - Release energy stored in organic compounds
- Photosynthesis
 - $\text{CO}_2 + \text{H}_2\text{O} \rightarrow \text{CH}_2\text{O} + \text{O}_2$
 - Sunlight provides energy
- Anoxygenic Photosynthesis
 - Use H₂ or H₂S instead of H₂O

Redox Balance

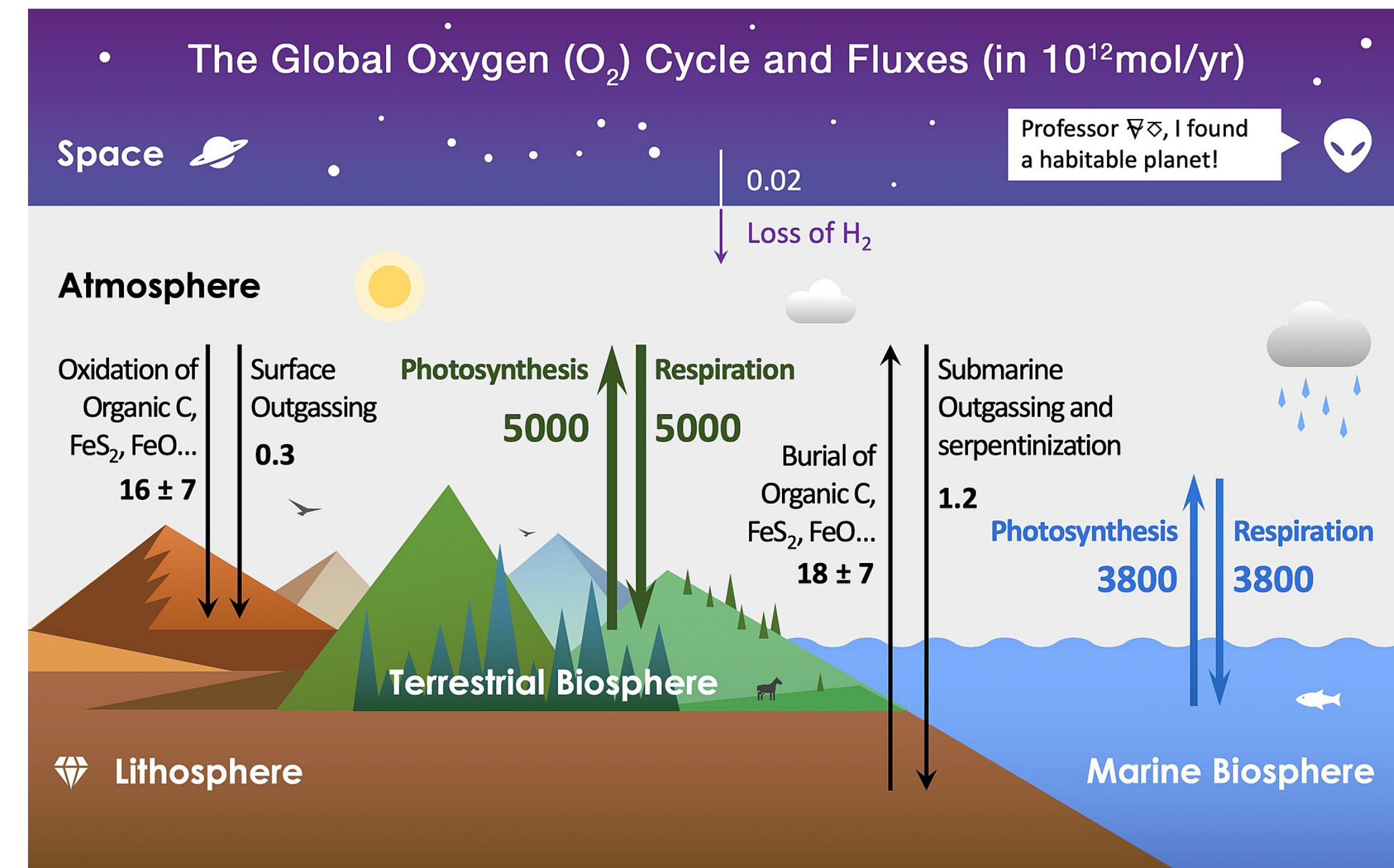
The Concept & Global Redox Budget

- Redox Balance = Conservation of Electrons
- Oxidation of one requires reduction of another
- Global Redox Budget = \sum (Oxidation Fluxes) – \sum (Reduction Fluxes)

Modern Oxygen Cycle

The Biological O₂ Cycle

- Source: oxygenic photosynthesis
 - Land photosynthesis: 60 Gt C yr⁻¹, 5×10^{15} mol O₂ yr⁻¹
 - Marine photosynthesis: 45 Gt C yr⁻¹, 3.8×10^{15} mol O₂ yr⁻¹
- Sink: respiration & decay (nearly balanced)
- Atmospheric O₂: 3.6×10^{19} mol
 - Residence time: $3.6 \times 10^{19} / 5 \times 10^{15} \approx 7000$ yr
 - (Each O₂ cycle once every 7000 yr)
 - (CO₂ residence time: 10 yr, preindustrial)



Modern Oxygen Cycle

The Geological O₂ Cycle: Source for O₂

Source process	Effect on O ₂ (×10 ¹² mol/yr)	Notes
Organic C burial	10 ± 3.3	Burial of reduced carbon in marine sediments (≈0.25% of marine NPP)
Pyrite (FeS ₂) burial	7 ± 3.6	Produced by bacterial sulfate reduction
FeO burial	0.9 ± 0.4	Sedimentary FeO contributes
H escape	0.024	Loss of H to space → net oxidation.
Total	≈ 18 × 10 ¹² mol/yr	—

- $\text{CH}_2\text{O} + \text{O}_2 \rightarrow \text{CO}_2 + \text{H}_2\text{O}$
 - $4\text{FeS}_2 + 15\text{O}_2 + 14\text{H}_2\text{O} \rightarrow 4\text{Fe}(\text{OH})_3 + 8\text{H}_2\text{SO}_4$
 - $4\text{FeO} + \text{O}_2 \rightarrow 2\text{Fe}_2\text{O}_3$
 - $2\text{H}_2\text{O} \xrightarrow{h\nu} \text{O}_2 + 4\text{H} \uparrow$
- Understand: reduced species are buried, releasing O₂ equivalents

Modern Oxygen Cycle

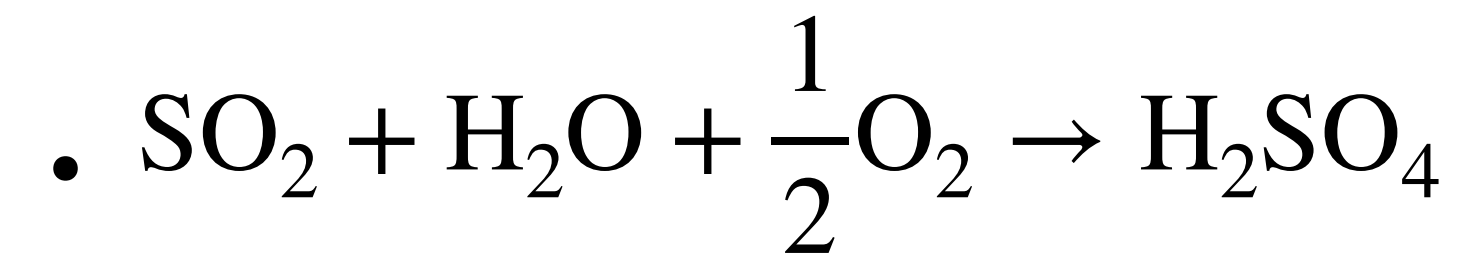
The Geological O₂ Cycle: Sinks for O₂

Sink process	Effect on O ₂ (×10 ¹² mol/yr)	Notes
Oxidation of organic C during weathering	7.5 ± 2.5	Continental weathering
Oxidation of FeS ₂	7.0 ± 3.6	Produces H ₂ SO ₄ → consumes O ₂
Oxidation of FeO	1.0 ± 0.6	Minor contribution
Surface volcanic gases (H ₂ , SO ₂ , H ₂ S)	0.3	Converted to sulfate via oxidation
Submarine H ₂ S oxidation	1.0	Hydrothermal vent processes
Serpentinization (Fe oxidation)	0.2	H ₂ generation at mid-ocean ridges
Total	≈ 17 × 10 ¹² mol/yr	nearly balanced with production.

Modern Oxygen Cycle

The Geological O₂ Cycle: Volcanic and Seafloor O₂ Sinks

- Volcanic gas oxidation

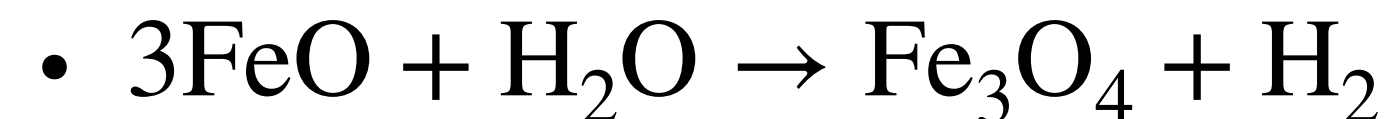


- Subaerial volcanism: $\sim 1 \times 10^{11}$ mol H₂/yr

- Mid-ocean ridge fluxes dominate ($\sim 80\%$ of total)

- Combined net “volcanic” O₂ sink $\approx 1.5 \times 10^{12}$ mol H₂/yr

- Serpentinization (蛇纹石化) :



- Reaction of ultramafic(超镁铁质) rocks(olivine橄榄石, pyroxene辉石) with hot sea water

- Generates H₂, consumes O₂ equivalents

- Minor today, possibly major in early Earth

Modern Oxygen Cycle

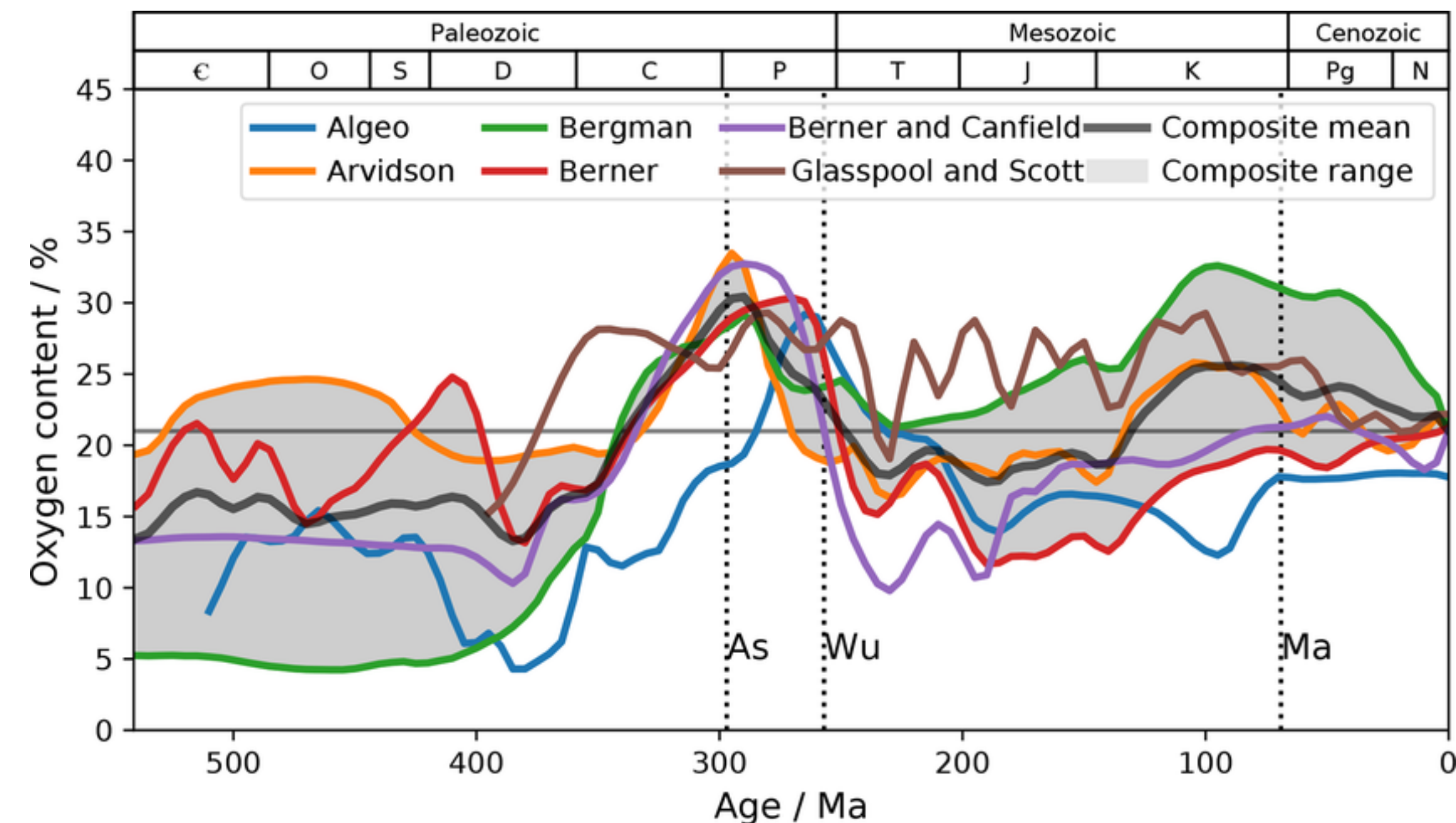
The Geological O₂ Cycle: Balance and O₂ Lifetime

- Net budget:
 - O₂ production: 17.9×10^{12} mol/yr
 - O₂ consumption: 17.0×10^{12} mol/yr
 - Effectively balanced within uncertainties($\pm 50\%$)
- Geologic lifetime of O₂
 - $\tau_{O_2} = \frac{3.6 \times 10^{19} \text{ mol}}{1.8 \times 10^{13} \text{ mol/yr}} \approx 2 \times 10^6 \text{ years}$
 - ≈ 2 million years
 - Timescale for large O₂ variations if sources/sinks shift
 - **Atmospheric O₂ is geologically stable**

Modern Oxygen Cycle

O₂ over Phanerozoic Time

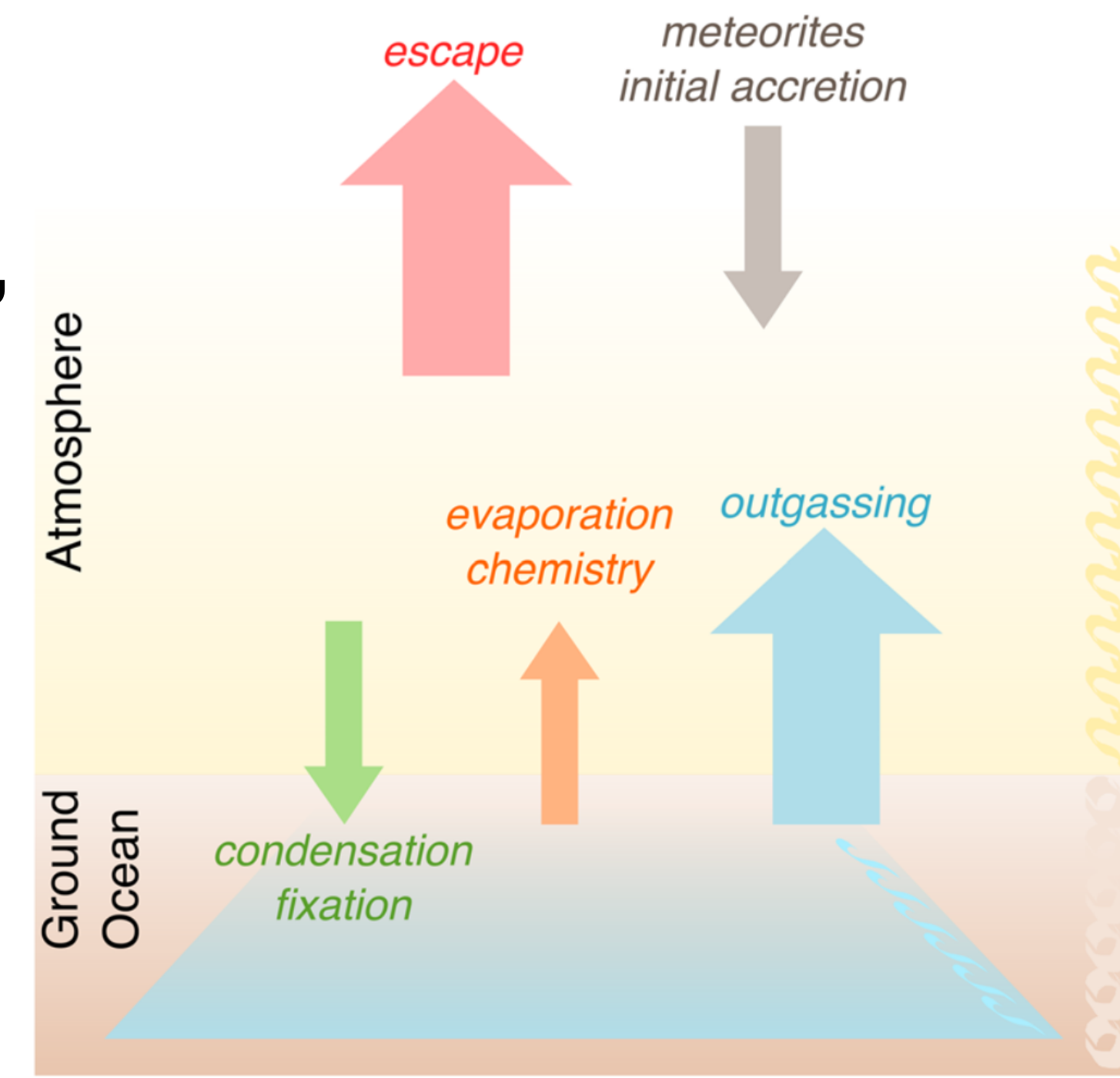
- High O₂ overall: Animal persistence implies O₂ never < ~ 10 % of present
- Most of time O₂ stayed ~ ± 50 % of present: fires give the upper bound
- Stabilizing feedbacks
 - weathering depends on O₂
 - pyrite burial inversely depends in O₂
 - newest sediment weather first



Modern Oxygen Cycle

Loss of Hydrogen to Space: Concept & Redox Effect

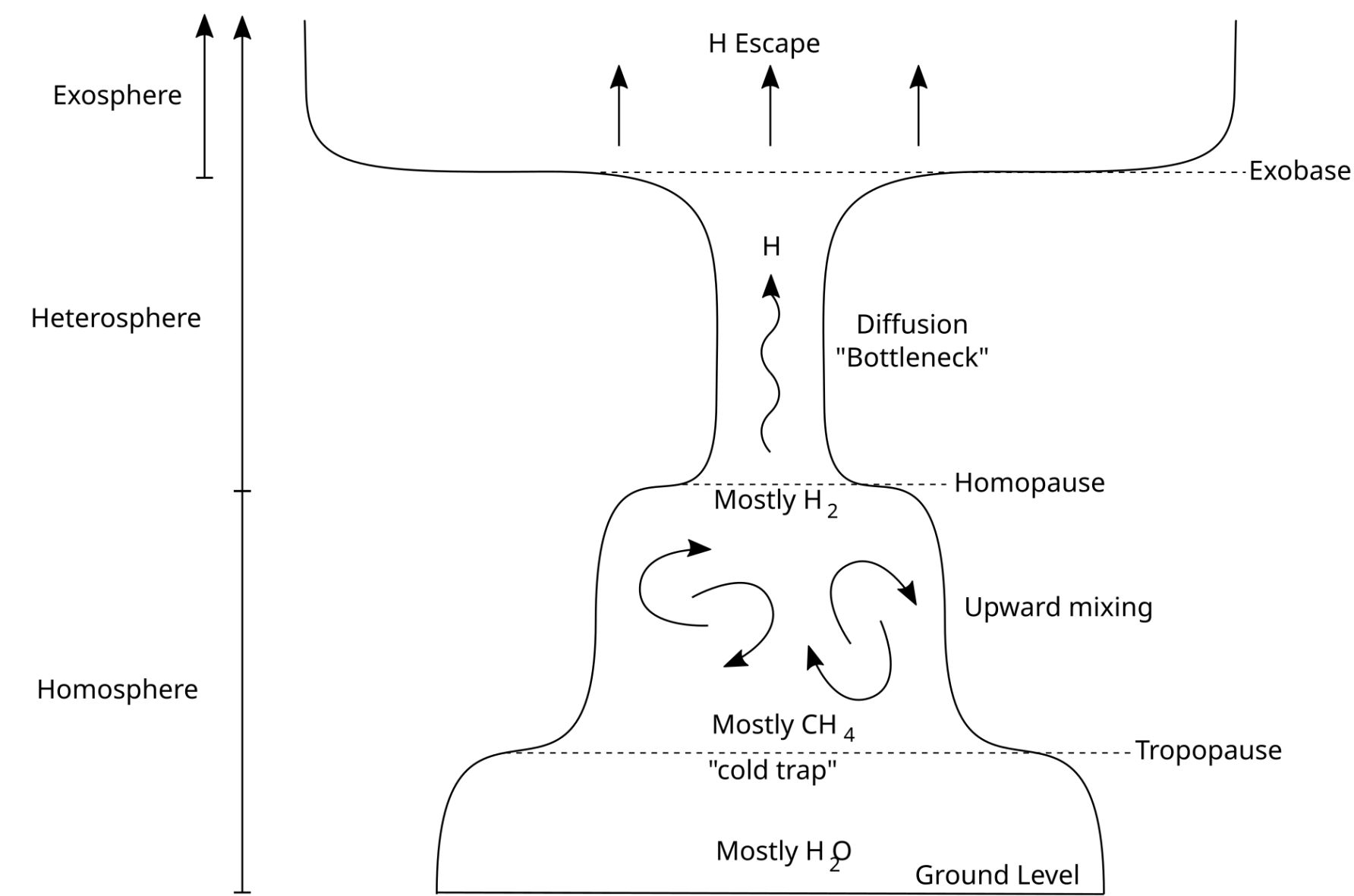
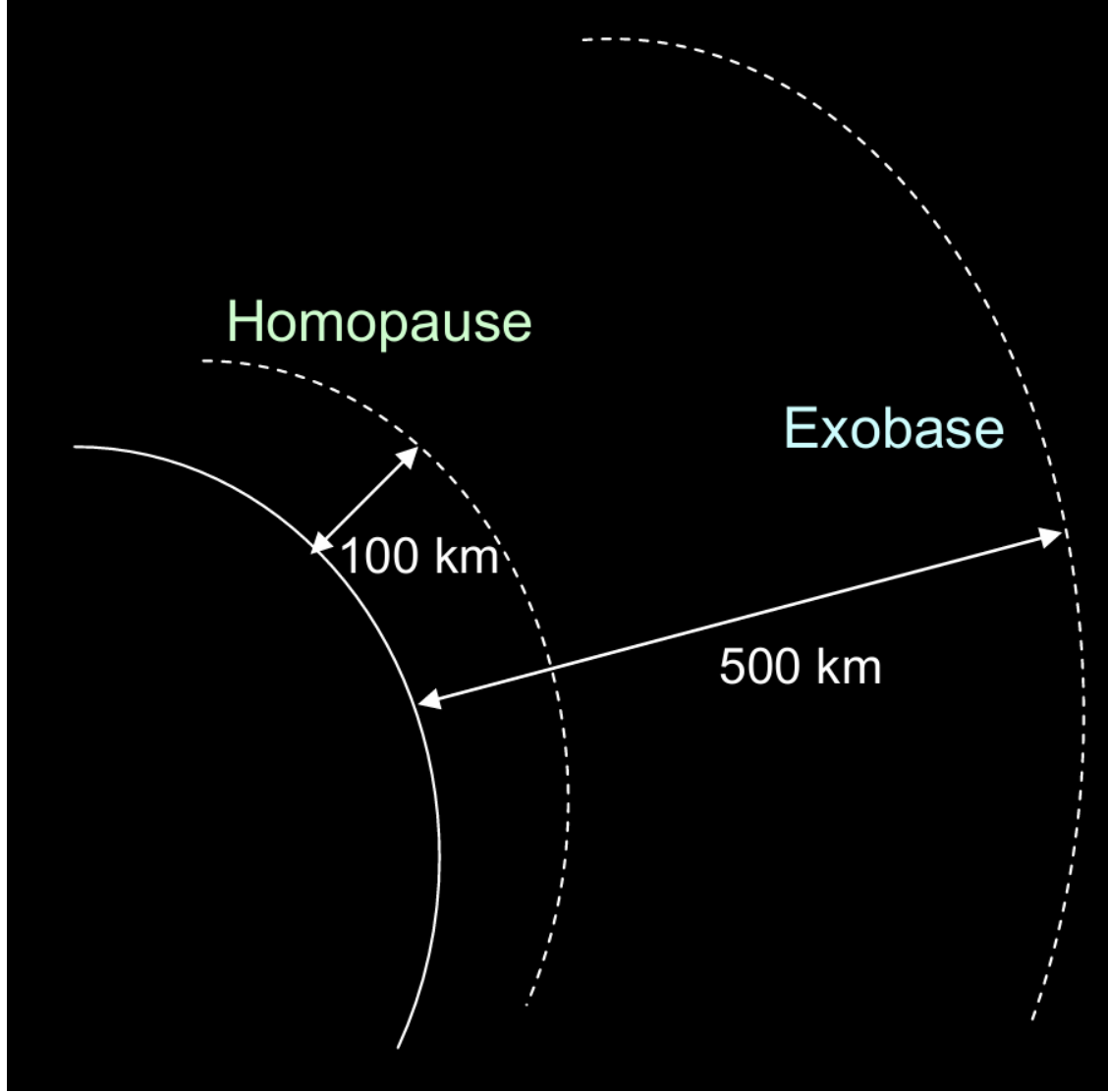
- Global Redox Budget:
 - Internal O_2 sources/sinks
 - Top boundary: Hydrogen escape to space
- Most H on Earth is H_2O , the loss of H leaves O_2 behind, producing a net oxidation of the planet



Modern Oxygen Cycle

Loss of Hydrogen to Space: Mechanism

- Escape occurs above the Exobase
 - (by thermal & non-thermal mechanisms)
- On Modern Earth, the escape rate is limited by diffusion through homopause



PLANET	KEY GASES LOST	DOMINANT MECHANISMS
Earth	Hydrogen	Charge exchange, Jeans, polar wind
	Helium	Polar wind, charge exchange
Early Earth	Hydrogen and moderately light gases, including neon	Hydrodynamic escape and drag
Venus	Hydrogen, helium	Charge exchange, sputtering
Early Venus	Hydrogen and moderately light gases, including oxygen	Hydrodynamic escape and drag
Mars	Hydrogen	Jeans
	Carbon, oxygen, nitrogen, argon	Sputtering, photochemical
Early Mars	All gases	Impact erosion
	Hydrogen and many heavier gases, including carbon dioxide	Hydrodynamic escape and drag
Early Callisto, Ganymede, and Europa	All gases	Impact erosion, hydrodynamic escape and drag
Titan	Hydrogen	Jeans, photochemical
	Methane, Nitrogen	Photochemical, (hydrodynamic flow?), sputtering
Early Titan	Hydrogen, methane, nitrogen	Hydrodynamic escape and drag
Pluto	Hydrogen, methane, nitrogen	Hydrodynamic escape?
HD 209458b and similar 'Hot Jupiters'	Hydrogen and light gases, including carbon and oxygen atoms	Hydrodynamic escape and drag

Modern Oxygen Cycle

Loss of Hydrogen to Space: Mechanism of Thermal Escape

- Neutral constituents in the upper atmosphere are in local thermodynamic equilibrium(LTE, 局部热力学平衡). Maxwellian function:

- $$f(\vec{x}, \vec{v}) = N \left(\frac{m}{2\pi kT} \right)^{3/2} e^{-\frac{mv^2}{2kT}} = N \left(\frac{1}{u_i \sqrt{\pi}} \right)^3 e^{-v^2/u_i^2}$$

- $$u_i = \sqrt{\frac{2kT}{m_i}}, \text{ thermal speed for species}$$

- Flux of Jeans Escape:

- $$\Phi_i(\text{escape}) = \int_0^{2\pi} \int_0^{\pi/2} \int_{v_{\text{esc}}}^{\infty} v_i \cos(\theta) f(v_i) v_i^2 \sin(\theta) dv_i d\theta d\Psi$$
$$= N_i \left(\frac{kT_e}{2\pi m_i} \right)^{1/2} \left(1 + \frac{m_i v_{\text{esc}}^2}{2kT_e} \right) e^{-\frac{m_i v_{\text{esc}}^2}{2kT_e}} = N_i \left(\frac{u_i}{2\sqrt{\pi}} \right) \left(1 + \frac{v_{\text{esc}}^2}{u_i^2} \right) e^{-\frac{v_{\text{esc}}^2}{u_i^2}}$$

- use the values of $v_{\text{esc}} = \sqrt{2GM/r}$ and u_i at the exobase (collisionless).

Modern Oxygen Cycle

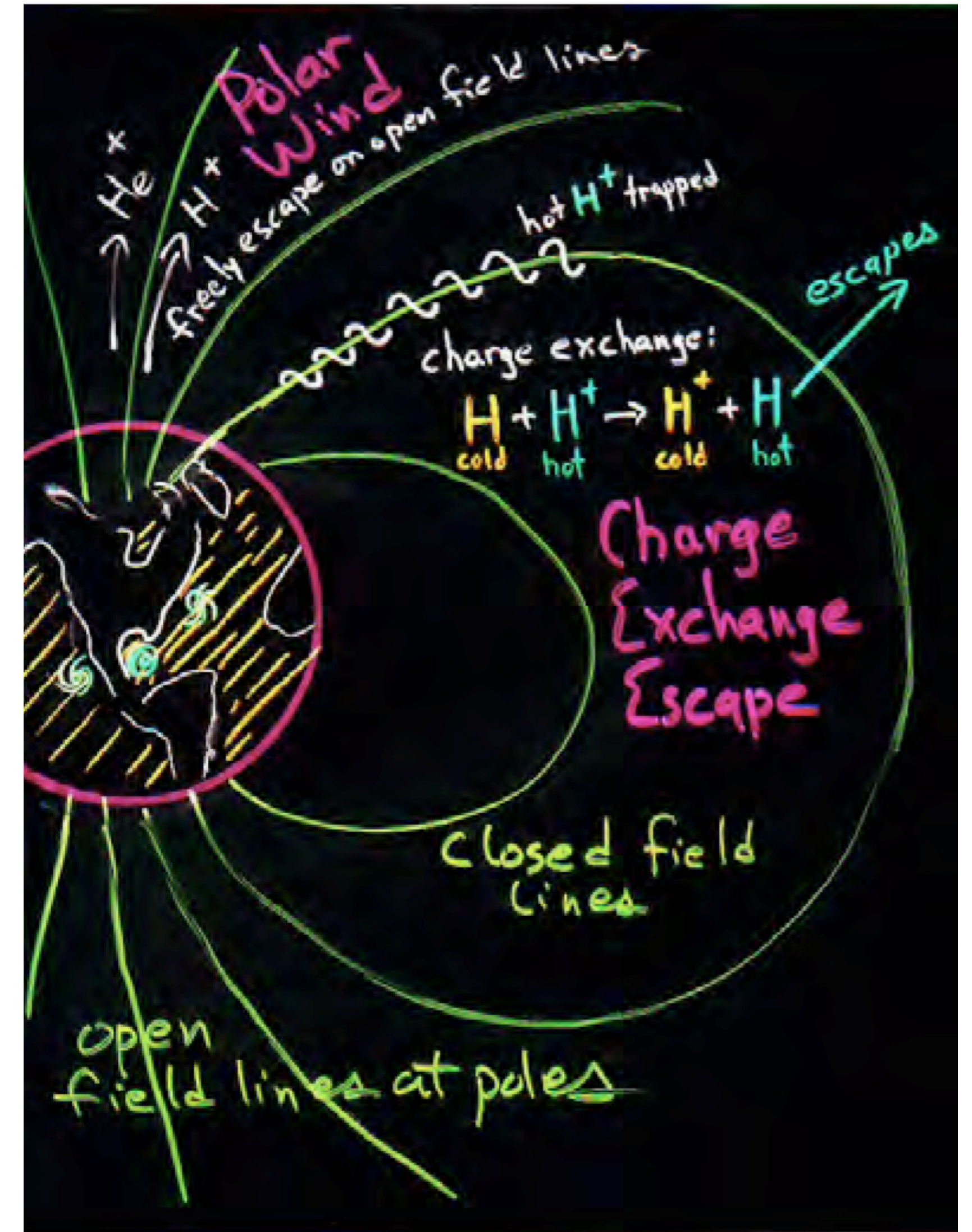
Loss of Hydrogen to Space: Mechanism of Thermal Escape

- We define Jeans parameter: $\lambda_{\text{ex}} = \frac{GMm_i/r}{kT} = \frac{v_{\text{esc}}^2}{u_i^2}$
 - (Gravitational Potential Energy/Thermal Energy)
- $\Phi_i(\text{escape}) = N_i \left(\frac{u_i}{2\sqrt{\pi}} \right) (1 + \lambda_{\text{ex}}) e^{-\lambda_{\text{ex}}}$
 - $\lambda_{\text{ex}} \uparrow$, $\Phi_i(\text{escape}) \downarrow$ exponential decay
 - Light particles' λ_{ex} is small, easy to escape (H & O ?)

Modern Oxygen Cycle

Loss of Hydrogen to Space: Mechanism

- Hydrodynamic escape can be regarded as an extreme form of Jeans escape, where the bulk flow carries gas away.
- For non-thermal escape, particles also need to exceed the local escape velocity v_{esc} .
- Since lighter species have higher thermal velocities at a given temperature, they are more likely to reach or exceed v_{esc} , and hence are more easily lost to space.



Modern Oxygen Cycle

Loss of Hydrogen to Space: Diffusion Limited Escape

- If the escape rate is efficient at exobase, the escape process is “limited” by the diffusion through homopause
- The vertical flux of a molecule is $\Phi_i = \Phi_i^{\text{mol}} + \Phi_i^{\text{eddy}}$
 - (molecular diffusion+eddy diffusion)
- Above the homopause, we only consider Φ_i^{mol} : (a light gas 1 diffuse through background gas 2)

$$\Phi_1^{\text{mol}} = -D_{12} n_1 \left(\frac{n^2}{n_1 n_2} \frac{df_1}{dz} + \frac{(m_2 - m_1)g}{kT} + \frac{\alpha_T}{T} \frac{dT}{dz} \right)$$

- (mixing ratio gradient+gravitational separation+thermal diffusion)
- D_{12} is binary diffusion coefficient, $f_1 = n_1/n$ is mixing ratio
- Under steady state condition approximation: $\Phi_l = D_{12} n_1 \frac{(m_2 - m_1)g}{kT}$
- $D_{12} = b_{12}/n$, where b_{12} is binary diffusion parameter, $\Phi_l = \frac{b_{12}(m_2 - m_1)g}{kT} f_1$
- In the end, we get: $\Phi_{\text{esc}} \approx \frac{b f_{\text{tot}}(H)}{H_a}$

Modern Oxygen Cycle

Loss of Hydrogen to Space: Diffusion Limited Escape

- $\Phi_{\text{esc}}(\text{H}_2) \approx 2.5 \times 10^{13} f_{\text{tot}}(\text{H}_2) \text{ molec cm}^{-2}\text{s}^{-1}$
 - $f_{\text{tot}}(\text{H}_2) = f(\text{H}_2) + \frac{1}{2}f(\text{H}) + f(\text{H}_2\text{O}) + 2f(\text{CH}_4) + \dots$

- At the homopause: H_2O (~4ppm), CH_4 (~1.6ppm)

- $f_{\text{tot}}(\text{H}_2) \approx 7.2 \times 10^{-6}$

- $\Phi_{\text{esc}}(\text{H}_2) \approx 1.8 \times 10^8 \text{H}_2 \text{ molec cm}^{-2}\text{s}^{-1}$

- Converted to global flux: $\approx 4.8 \times 10^{10} \text{ mol H}_2 \text{ yr}^{-1}$

- Equivalent O_2 production: $\frac{1}{2} \Phi(\text{H}_2) \approx 2.4 \times 10^{10} \text{ mol O}_2 \text{ yr}^{-1}$

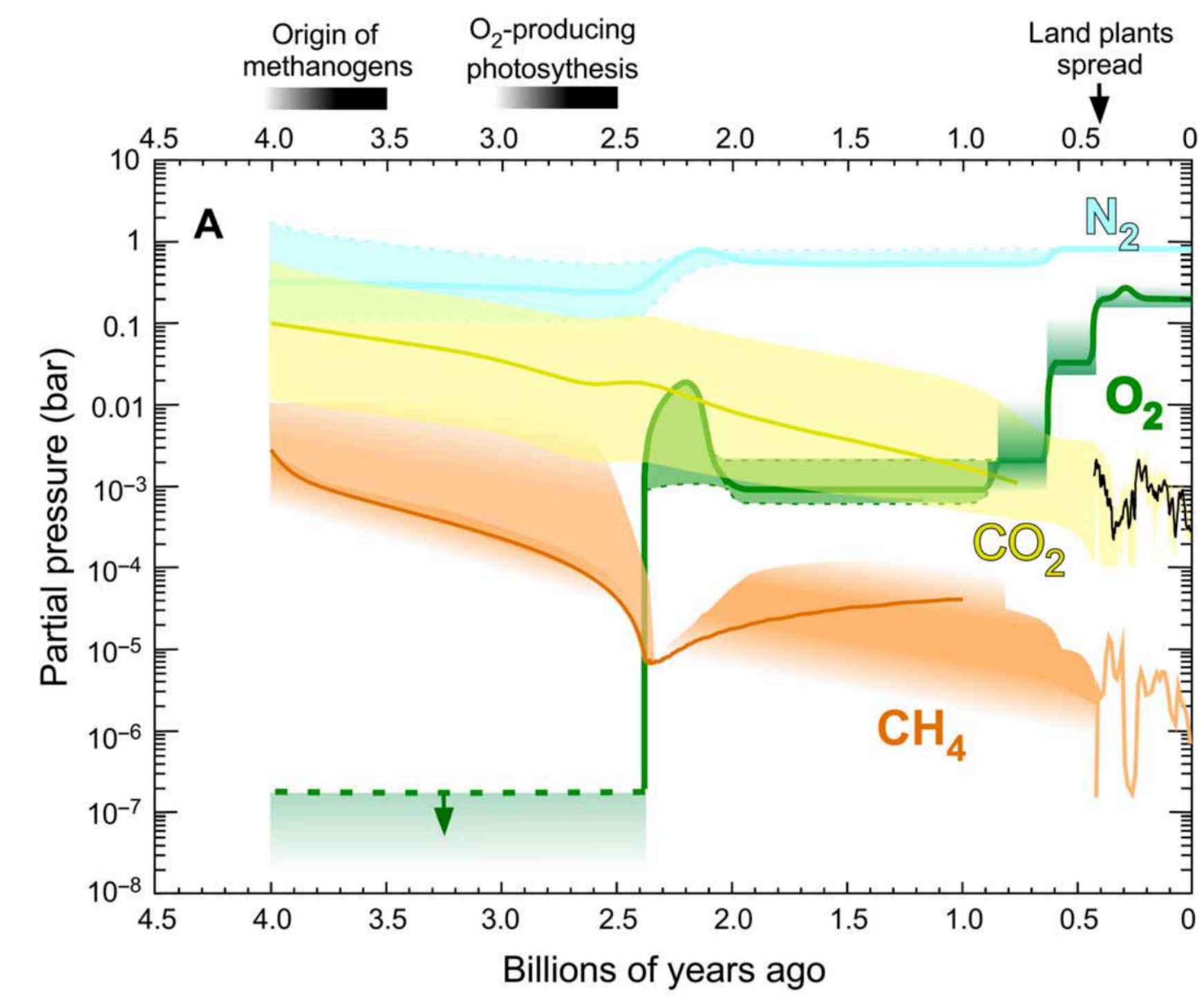
- Compare to geologic O_2 production: $\approx 1.8 \times 10^{13} \text{ mol yr}^{-1} \rightarrow \text{negligible (} \sim 0.1\%)$

Modern: O_2 -rich atmosphere prevents H accumulation
Early Earth???

Early Earth Oxygen Cycle

Overview & Timing of O₂ Rise

- Before the GOE, atmospheric O₂ was very low
 - (MIF-S, 硫同位素质量无关分馏)
- Oxygenic photosynthesis likely invented by 2.7-2.5Ga, but O₂ stayed low
- Why the delay between photosynthesis and rise of O₂?



Early Earth Oxygen Cycle

Evidence & Biosignatures

- Evidence for early oxygenic photosynthesis:
 - Mo, U, N isotopes in black shales(页岩)
 - Biomarkers(2-methylhopanes, steranes)
 - Microfossils / stromatolites(叠层石)
- Overall data support cyanobacterial evolution before GOE
- ? Biology was ready, but environmental sinks kept O₂ low



Early Earth Oxygen Cycle

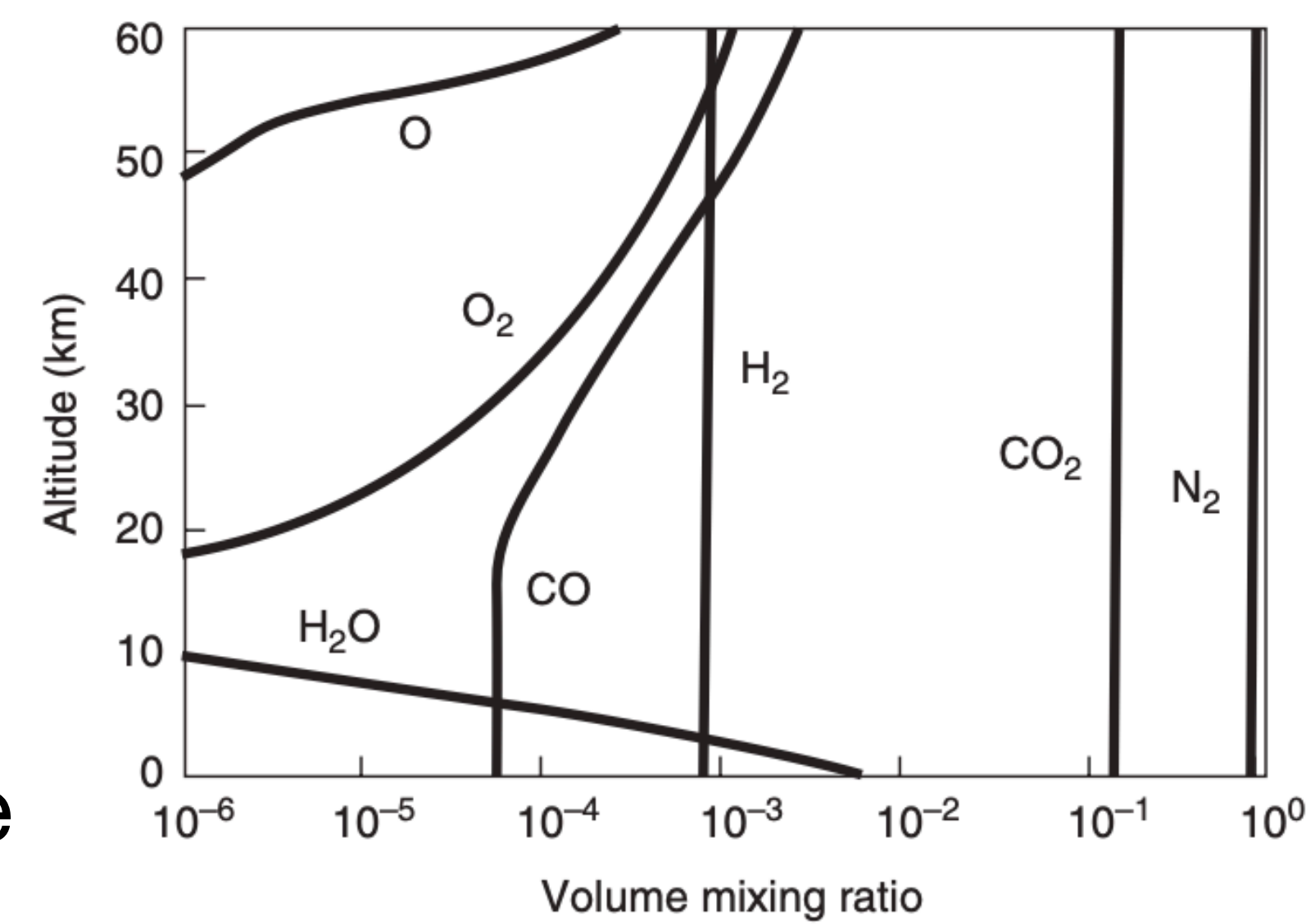
Prebiotic / Very-low-O₂ Redox Budget

- Consider the abiotic Earth (no organic burial, negligible abiotic organic production)
- With low oceanic sulfate in the Archean, H₂S formation at ridges was limited, volcanic gases became dominant O₂ sinks (surface + submarine outgassing)
- Sources of O₂ prebiotically are few: effectively only H escape to space (from H₂O photolysis)
 - Modern-style calculation gives O₂ eq. $\sim 1.3 \times 10^{10}$ mol/yr
 - (from H₂O-sourced H escape)
- Net: sinks \gg sources \rightarrow O₂ remains trace

Early Earth Oxygen Cycle

H₂ escape controls & expected H₂, O₂ levels

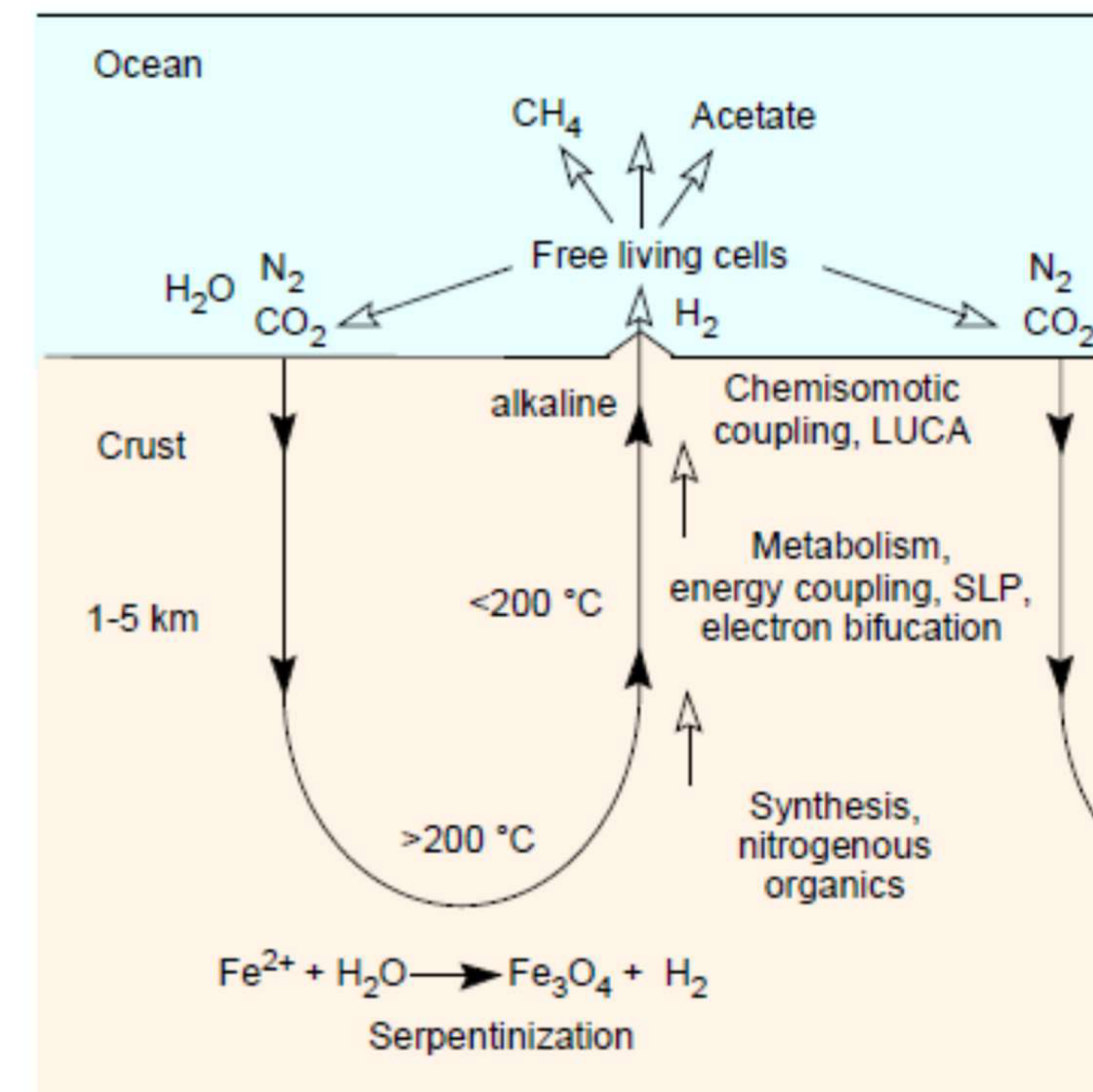
- H escape on early Earth is diffusion-limited at the homopause
 - (flux \propto total H mixing ratio)
- Photochemical models predict H₂ mixing ratio $\sim 10^{-3}$ (order-of-magnitude), consistent with enhanced H outgassing
 - (by higher Archean geothermal heat flow and serpentinization)
- O₂: a trace species
 - $\sim 10^{-3}$ in the upper stratosphere, but near zero at the surface
 - (rapidly consumed by H₂).



Early Earth Oxygen Cycle

Oxidation of Mantle & Crust (delaying GOE)

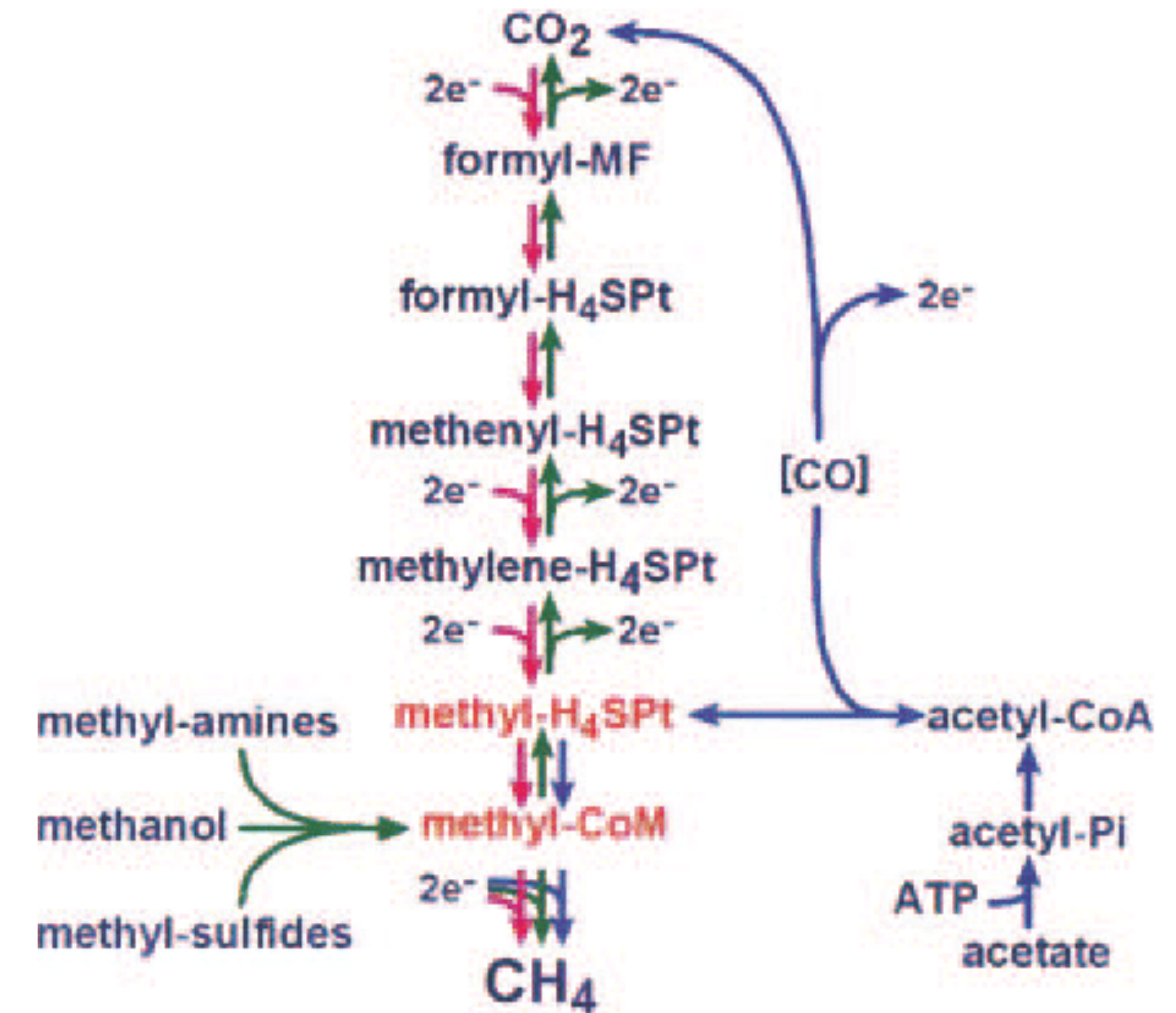
- An H_2 -rich atmosphere implies O must go somewhere
 - oxidation of crust and mantle can store O_2 -equivalents.
- Processes:
 - serpentinization (蛇纹石化, $\text{Fe}^{2+} \rightarrow \text{Fe}^{3+} + \text{H}_2$)
 - weathering and subduction of ferric iron
 - progressive crustal oxidation.
- Large O sinks in the solid Earth: delay atmospheric O_2 rise until these sinks diminish.



Early Earth Oxygen Cycle

Biotic Archean & Role of CH₄

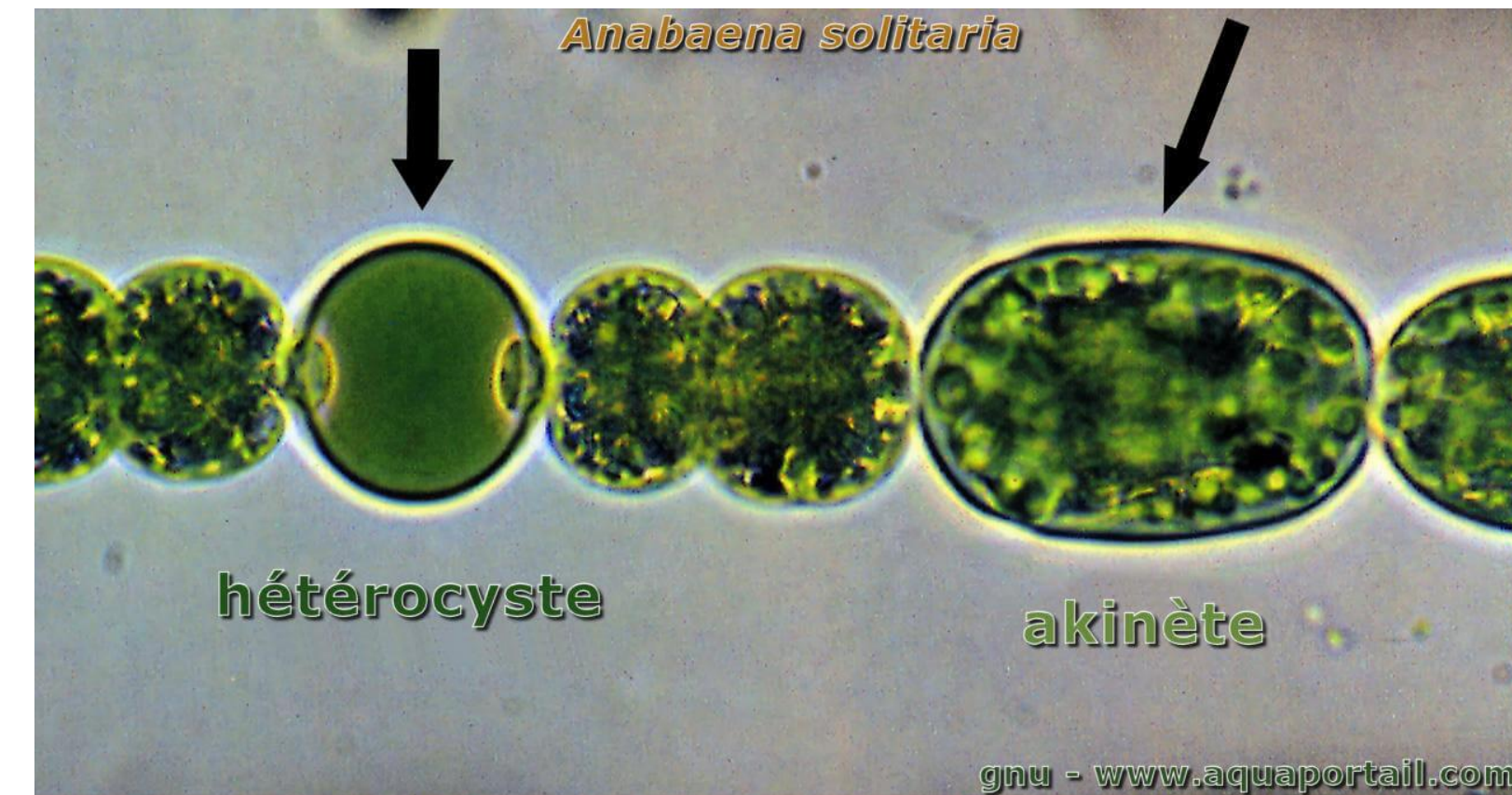
- After life becomes widespread
 - methanogens(甲烷菌): $\text{H}_2 + \text{organic C} \rightarrow \text{CH}_4$
- High CH₄:
 - greenhouse warming
 - affects redox budget



Rise of Atmospheric Oxygen

Why O₂ stayed low before GOE?

- O₂ producers (cyanobacteria) likely existed earlier, but O₂ remained low.
- Biological “brake”:
 - nitrogenase(固氮酶) is O₂-sensitive (anaerobic, 厌氧)
 - early cyanobacteria may have lacked protection, like heterocysts(异形胞)
 - limited spread
- Geochemical sinks dominated:
 - reduced volcanic/hydrothermal gases
 - crust & mantle oxidation
 - low sulfate → less pyrite burial (smaller O₂ source)
- Global redox budget net \approx zero or negative → O₂ cannot accumulate.



Rise of Atmospheric Oxygen

Why did O₂ rise at GOE?

- Biology improves:
 - cyanobacteria evolve nitrogenase protection
 - broader spread
- Sources strengthen / sinks weaken:
 - increased: S & C recycling, organic burial
 - declined: abiotic H₂ supply, hydrothermal(热液) sinks
 - seafloor shifts ultramafic(超镁铁质) → mafic(镁铁质), less reduced gas production
 - volcanism shifts to more subaerial
- Small, long-term positive redox imbalance → atmospheric O₂ rises
- Still uncertain, need improved tectonic–Earth-system models.....



Q&A